International Astrobiology Workshop@Sagamihara 2013.11.28.

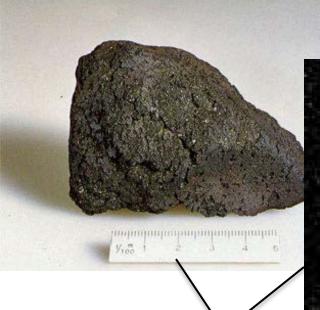
Evolution of Interstellar Organics to Meteoritic and Cometary Organics: Approaches by Laboratory Simulations

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Outline

- Origins of meteoritic organics and origins of life
- Possible formation and alteration of intertstellar complex amino acid precursors
- Cosmic dusts (IDPs) as carriers of organics to the Earth
- The Tanpopo Mission: Capture of cosmic dusts and exposure of organics in space
- Conclusion

Wide variety of extraterrestrial organics have been detected.



←Meteorites (Yamato791198)



← Comets (Wild 2)

IDPs ↓

✓ Complex organics
✓ Amino acids
✓ Organics seemed to be formed at low temperature

Origins of meteoritic amino acids

- Strecker synthesis in parent bodies of meteorites HCN + RCHO + NH₃ \rightarrow NH₂CHRCN $\xrightarrow{}_{H_2O}$ amino acids
- Fisher-Tropsch-type synthesis in solar nebula $CO + NH_3 + H_2O \longrightarrow \text{amino acids}$ $\Delta \text{ (metal catalysts)}$
- Alteration of interstellar complex organics



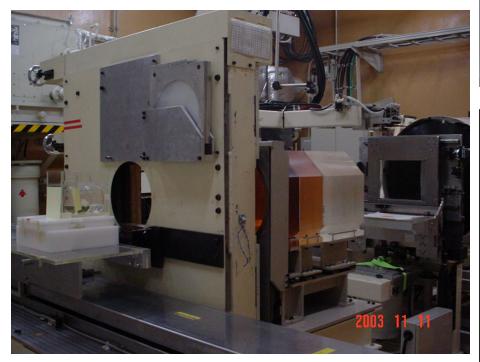




Particles Irradiation of Simulated Interstellar Media (Ice / Liquid / Gas)

Protons@Tandem Accelerator(TIT) \rightarrow

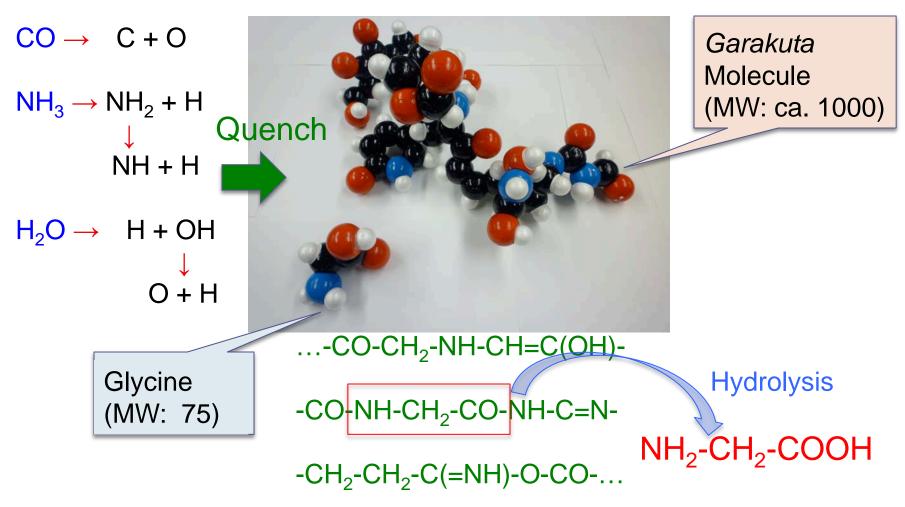
Heavy ions@HIMAC(NIRS) \downarrow





- Molecular weights: Some thousands
- In situ formation of complex (solid) organics
- Amino acids were yielded after hydrolysis
- Nucleic acid bases were formed

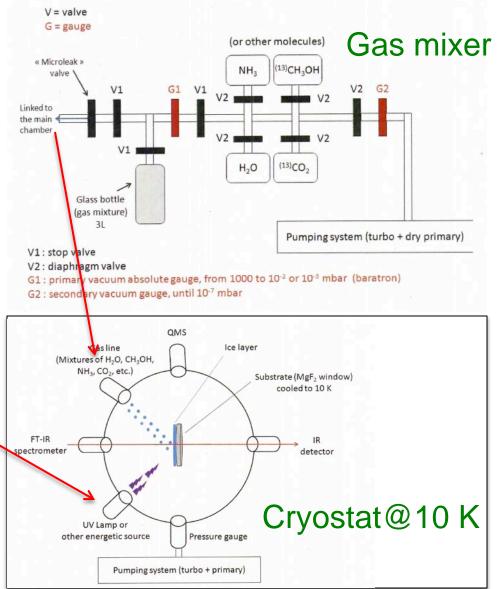
Formation of amino acids from complex organics (*Garakuta* molecules)

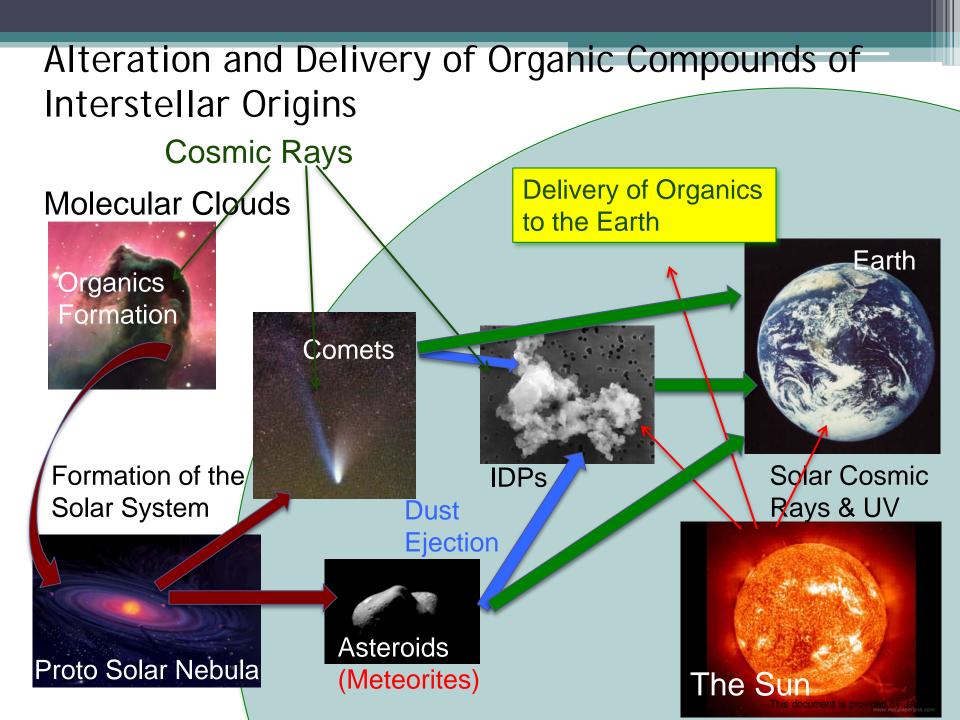


Heavy ion bombardent of Simulated Interstellar Ices

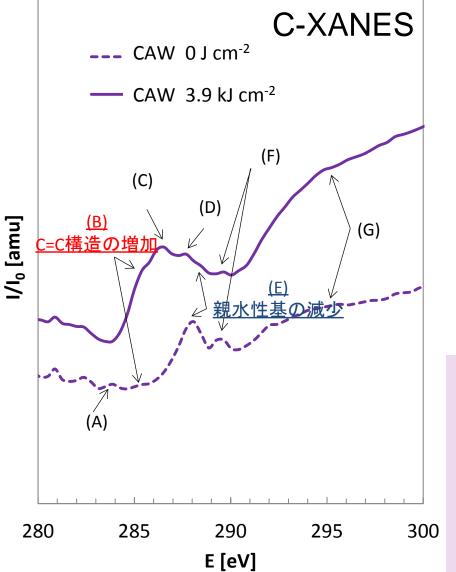








Alteration of Interstellar Organics (CAW) in Protosolar Nebula by Soft X-rays





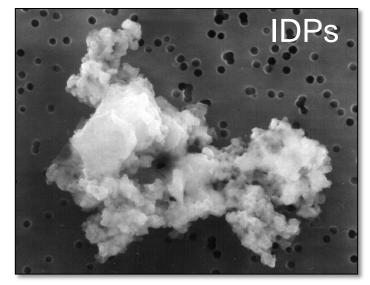
NewSUBARU, Univ. Hyogo □ Decrease of Hydrophilic C=O → (E) Amidyl

□ Increase of Hydrophobic C=C \rightarrow (B) Aryl, vinyl-keto

300 ⇒ Formation of Insoluble Organics This document is provided by JAXA.

Delivery of Organics to Earth: Meteorites vs. Cosmic Dusts (IDPs)

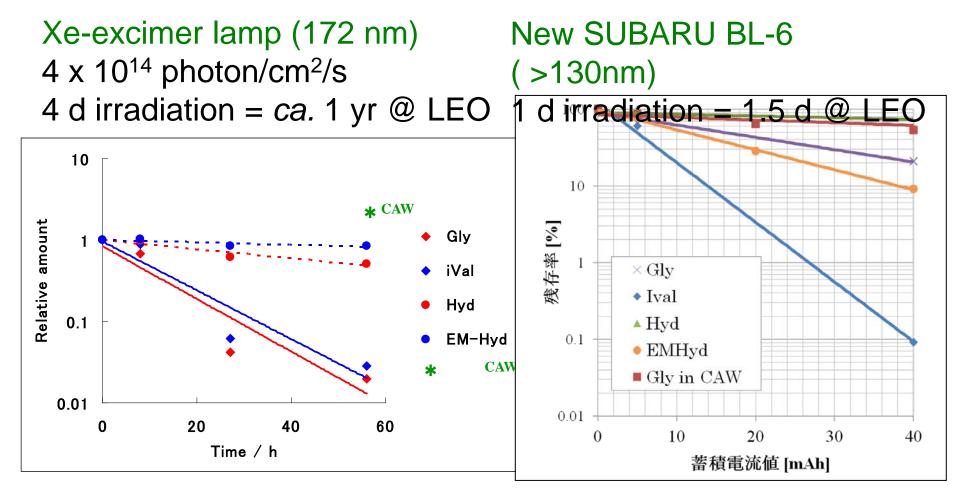
IDPs delivered more organics to to the Earth than meteorites and comets
IDPs delivered organics more safely than meteorite and comets.



 IDPs are directly exposed to solar / cosmic radiation.
IDPs are easily contaminated from terrestrial biosphere.



UV Irradiation of Amino Acids & Their Precursors

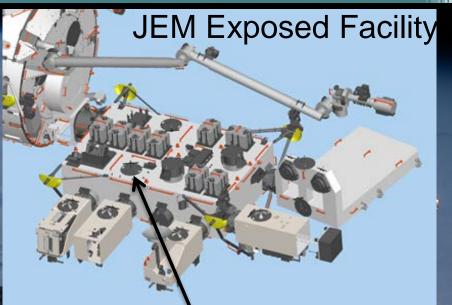


Amino acid precursors are much more stable than free amino acids against UV, X-rays, radiation and heat.

11

The International Space Station

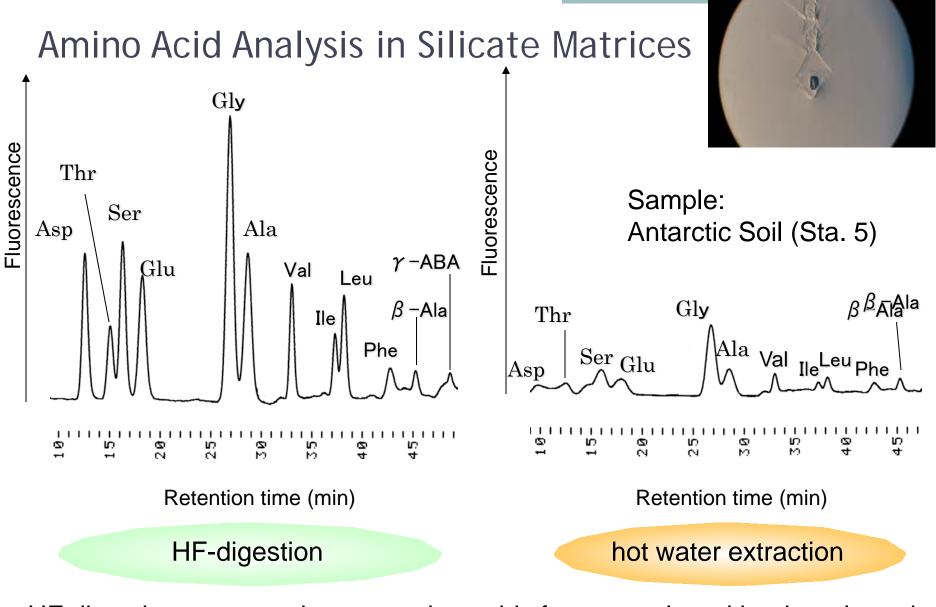
The Tanpopo Mission: Capture of Cosmic Dusts & Exposure of Organics





40 cm ExHAM

Aerogel



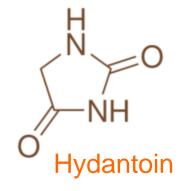
HF digestion gave much more amino acids from samples with mineral matrix

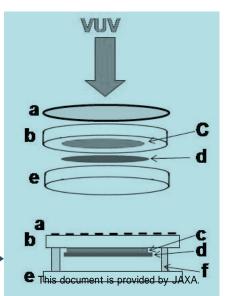
Exposure of Organic Compounds in the Tanpopo Mission

- Amino acids and their precursors
- •Glycine
- •Hydantoin (A precursor of glycine) \rightarrow
- •Isovaline (α -methyl non-protein amino acid)
- •5-Ethyl-5-methyl hydantoin
 - (A precursor of Isovaline)
- "CAW" (Complex amino acid precursors produced by proton irradiation of a mixture of CO, NH₃ and H₂O)

Alanine Thin Film (as a VUV dosimeter) \rightarrow







Summary

- Complex organic compunds with high molecular weights can be formed in simulated interstellar environments by high energy particles bombardment
- Meteoritic organics could have been formed by alteration of interstellar complex organics: Soft X-rays from the young Sun could be important energy source for it, as well as aqudous / hydrothermal alteration in parent bodies of meteorites.
- Major carriers of extraterrestrial organics could have been cosmic dusts (IDPs).
- Cosmic dusts will be collected in space with aerogel in the Tanpopo Mission
- Amino acids in cosmic dusts will be analyzed after HF digestion.

Thank you for your attention!

Acknowledgements

- Dr. Akihiko Yamagishi (Tokyo Univ. Pharm. Life Sci.) and the Tanpopo WG
- Dr. Issay Narumi (Toyo Univ.; γ -irradiation)
- Dr. Sunao Hasegawa (JAXA/ISAS; 2-stage light gas gun experiments)
- Dr. Ken Takayama (KEK; Digital accelerator)