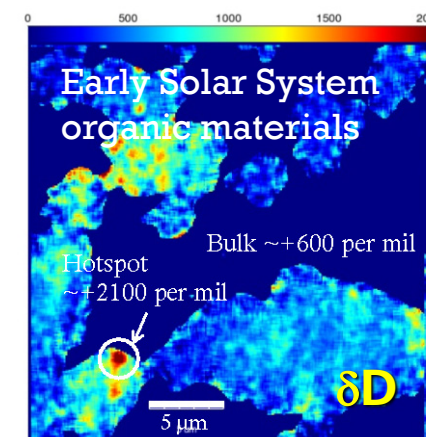
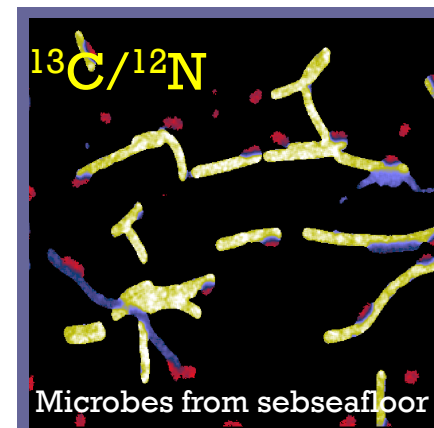
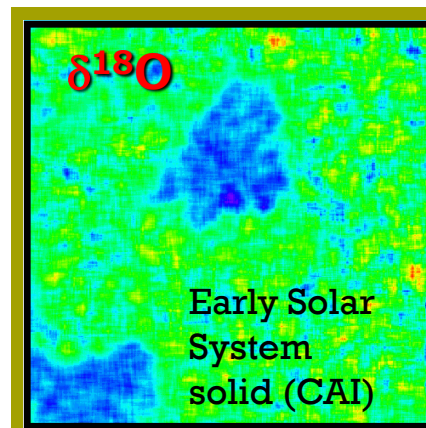




Ultra High Spatial Resolution Ion Imaging with a NanoSIMS Ion Microprobe: Applications to Astrobiology.

Motoo ITO

Kochi Institute for Core Sample research,
JAMSTEC



In collaboration with many scientists

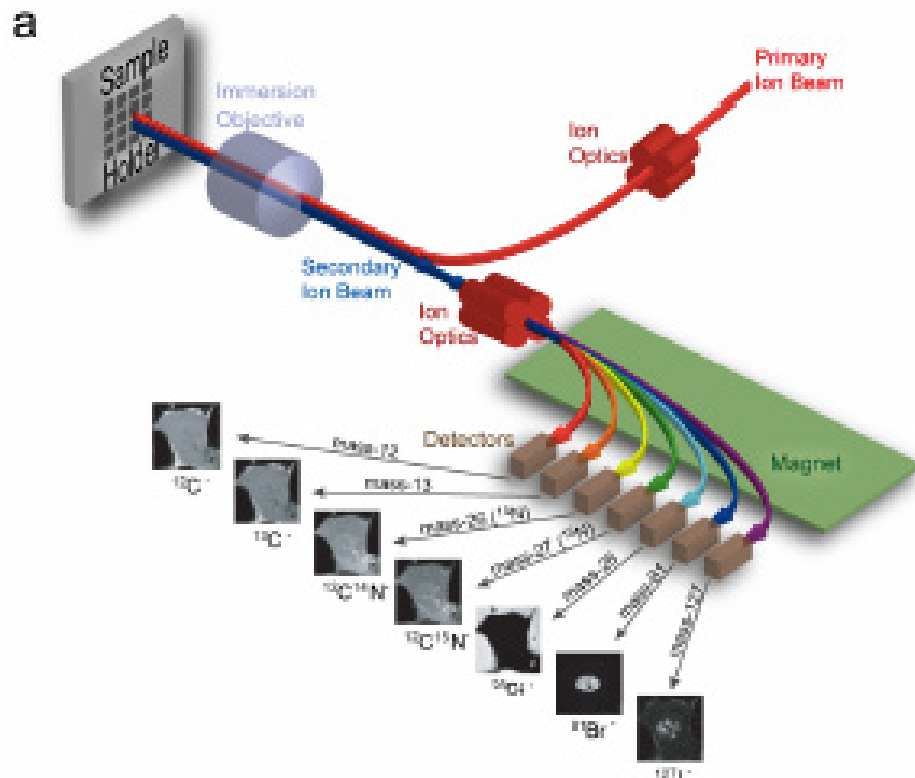
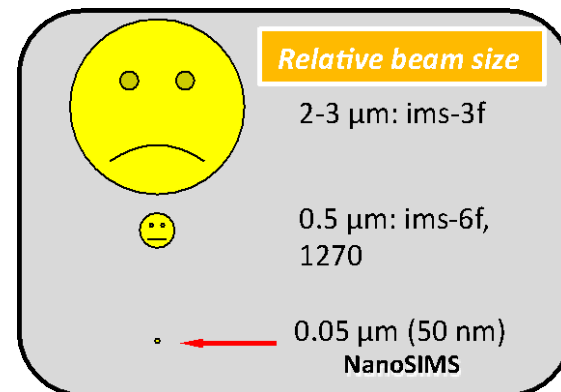
Microbiology: Y. Morono, F. Inagaki (JAMSTEC), T. Terada (MWJ)

Hayabusa Organic materials: M. Uesugi², H. Naraoka³, H. Yabuta⁴, F. Kitajima³, H. Mita⁵, Y. Takano¹, Y. Karouji², T. Yada², Y. Ishibashi², T. Okada² and M. Abe²

¹JAMSTEC. ²JAXA/ISAS. ³Kyushu University. ⁴Osaka University. ⁵Fukuoka Inst. Tech.



A very basic of NanoSIMS Ion Microprobe



Steinhauser et al. (2012) Nature

- Ultra high spatial resolution: ~50 nm for Cs, ~200 nm for Oxygen
- Up to 7 isotopes/elements can be measured simultaneously.
 - $M_{\max}/M_{\min} = 21$ (~25.4)
- Acquire isotope/elemental images
 - Clearly see 2-D distributions in sample
- Coordinated studies with TEM, SEM, and FIB systems are becoming routine.
- Lack of molecular info.

+ What we can do with NanoSIMS ion imaging

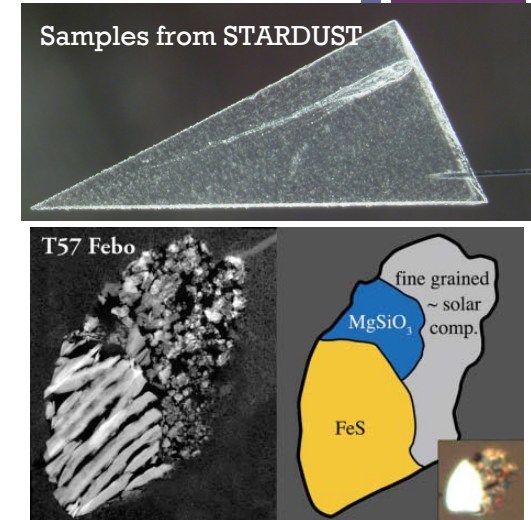
	Micro Analysis	Imaging
Cs primary ion source		
H (H, D)	N/A	~200 nm
C (^{12}C , ^{13}C)	N/A	~50 nm
O (^{16}O , ^{17}O , ^{18}O)	N/A	~50 nm
	3-5 μm	~400 nm
N (^{14}N , ^{15}N)	N/A	~50 nm
Volatile elements (H_2O , CO_2 , F, Cl, S etc.)		
High-precision O isotopic measurements (~0.01%)		
Oxygen primary ion source		
^{26}Al - ^{26}Mg : age dating	5-20 μm	0.2-1.5 μm
^{53}Mn - ^{53}Cr : age dating	3-5 μm	N/A
^{60}Fe - ^{60}Ni : age dating	~20 μm	N/A
Rare Earth Elements	8-15 μm	2 μm
Major and trace element abundances: Water in fluid inclusions, minerals and organic materials in ice		

On-going project

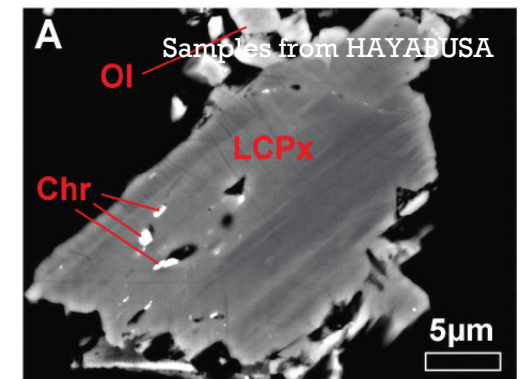
Oxygen ion source will be upgraded to get better spatial resolution.

+ Astrobiology Samples: Organics, microbes, minerals

- We are expecting to have samples by TANPOPO (PI: Prof. Yamagishi) and Hayabusa 2 those are similar to the samples by STARDUST and Hayabusa missions.
- Organics (carbonaceous materials)
 1. Very low temp. chemical processes in Molecular Cloud
 2. Origin of life, organics
- Minerals (micrometeorite, asteroidal particles)
 1. Presolar materials: origin of the Solar system
 2. Water bearing minerals: origin of Earth's water
 3. High temp. processes near the Early Sun
 4. Mineral-Water interactions during planetesimals
 5. Aqueous alteration, thermal and shock metamorphism



Brownlee et al. (2006)



Nakamura et al. (2011)



Our analytical techniques for organics, biology and mineral samples so far

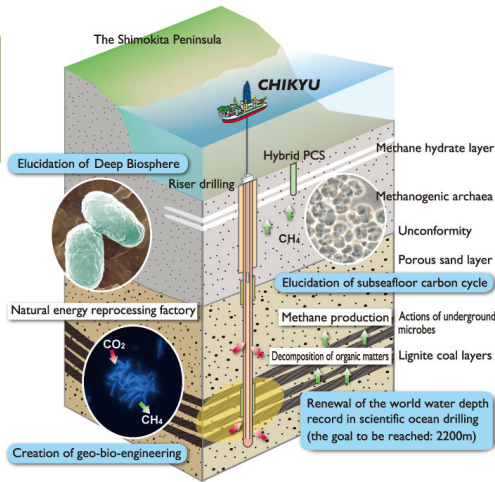
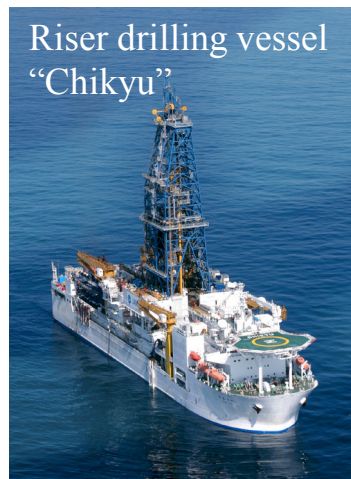


■ Biology sample

- Microbes from deep subsea floor

■ Organics

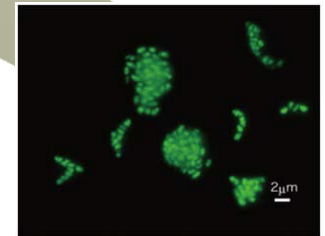
- IOMs from Carbonaceous chondrites,
- Hayabusa category 3 organic samples



Core sediments



Incubation with
 ^{13}C labeled
and/or ^{15}N
labeled sources
(e.g., glucose,
amino acid)

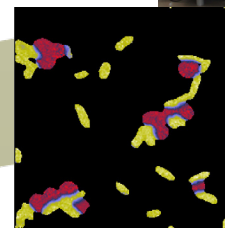
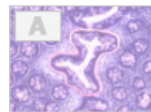


Deep subseafloor
microorganism

A sequence for biology sample: From deep subseafloor to Laboratory

I GOT IT!

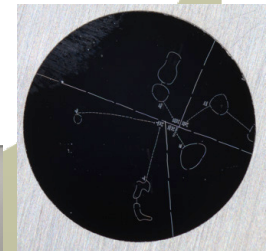
Extract DNA
from the
microbe that
contain ^{13}C
and/or ^{15}N
enrichments



Ion imaging



^{13}C labeled
Micro organisms



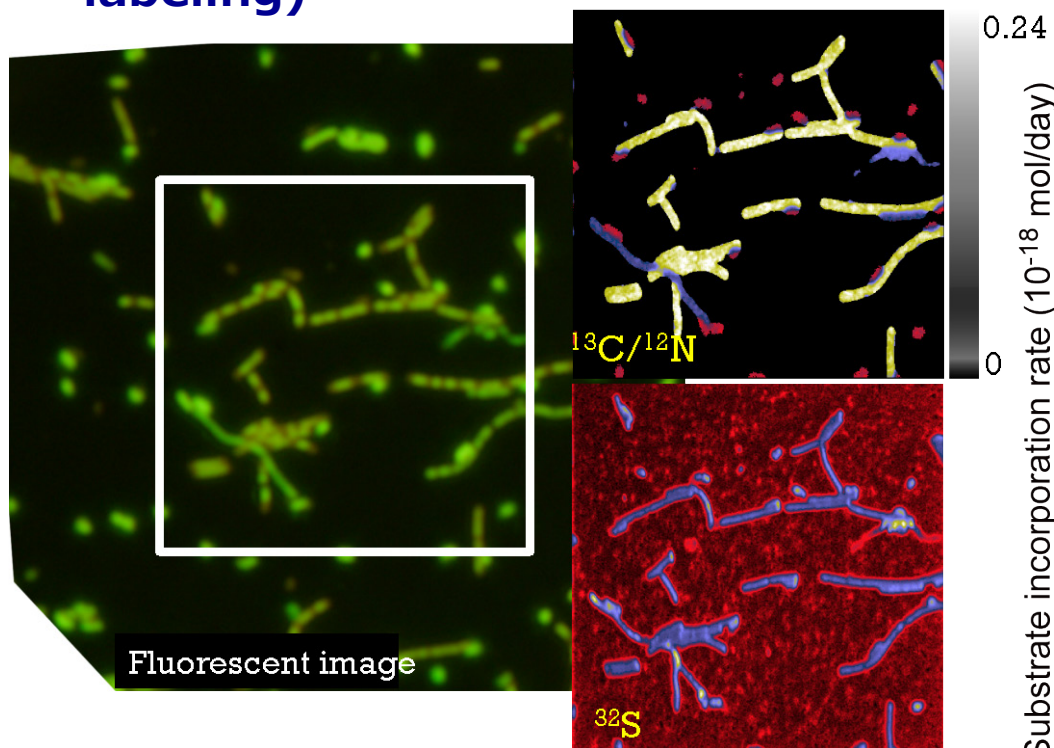
Microbes put
on Carbon
membrane
with ITO
coating

Laser Microdissection



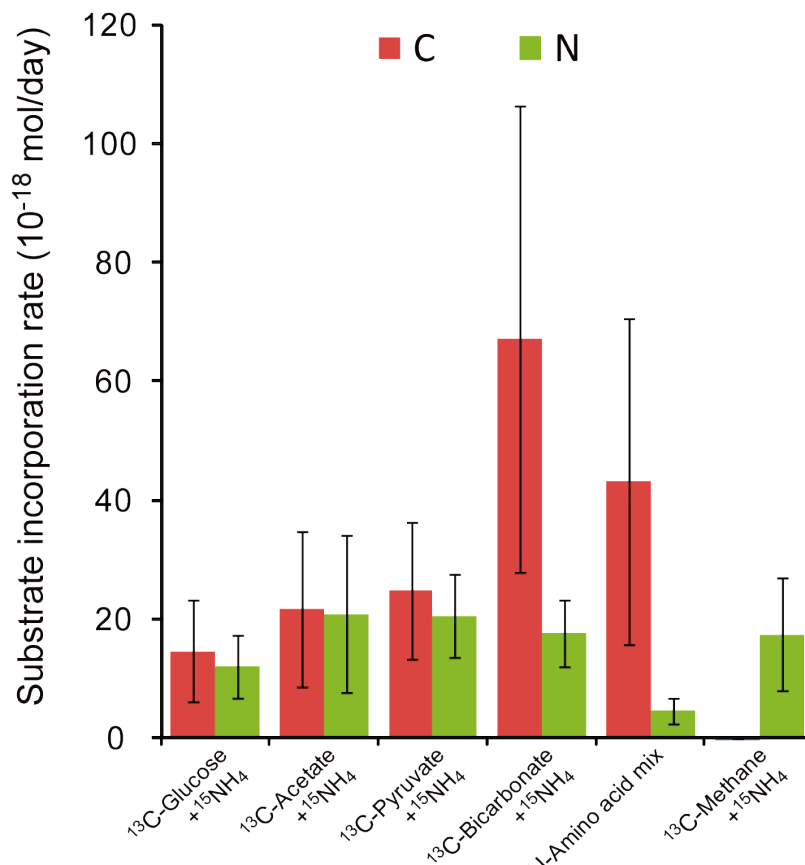
Isotope image for microbes from deep subseafloor

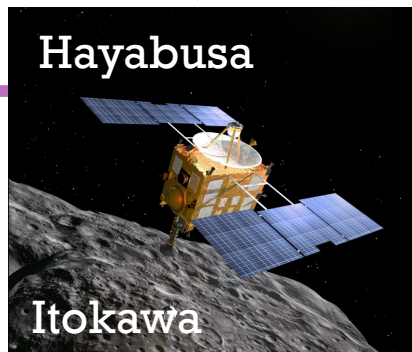
Stable isotope probing for activity measurement (e.g., ^{13}C , ^{15}N labeling)



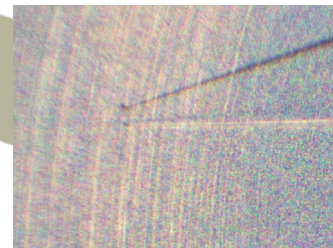
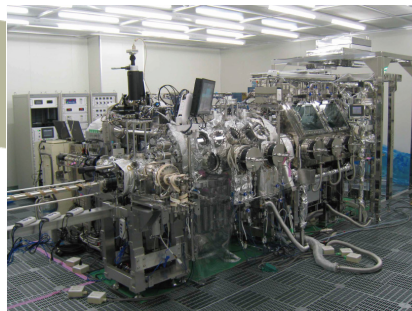
Deep seafloor microbes by the reactor enrichment culture using $^{13}\text{CO}_2$

Activity can be calculated for each substrate

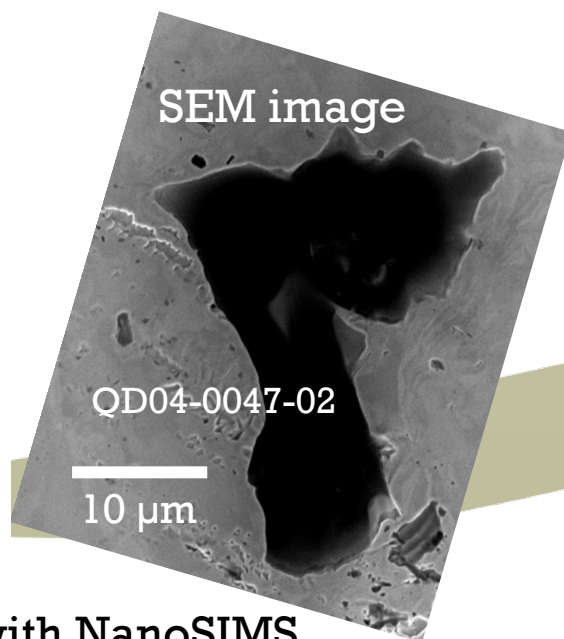
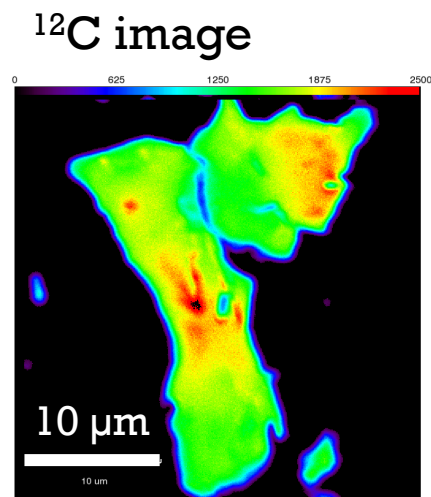




Sample curation @ JAXA

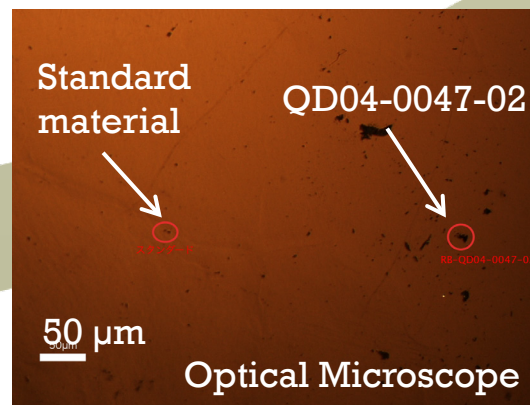
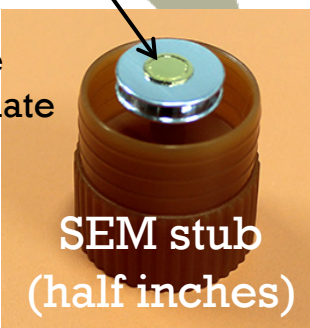


A sequence for Hayabusa Cat 3 organic sample



Samples were pressed on Au plate

Au plate



Isotope imaging with NanoSIMS

+ Isotopic anomalies in primitive solar system organic materials

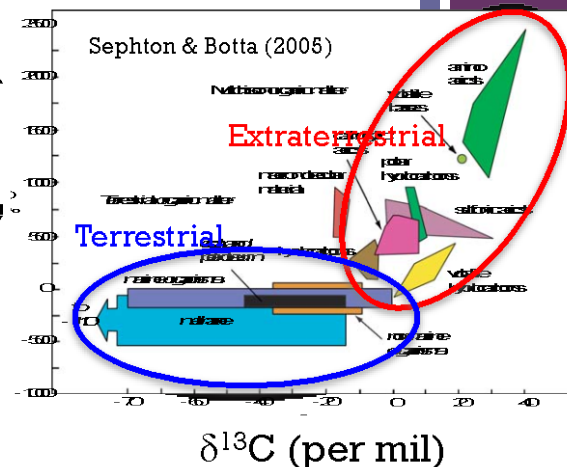
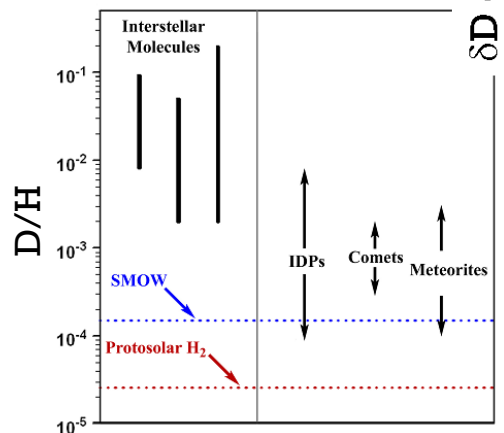
1. Cold interstellar clouds

- large δD values have been observed
- large $\delta^{15}N$ values have been predicted

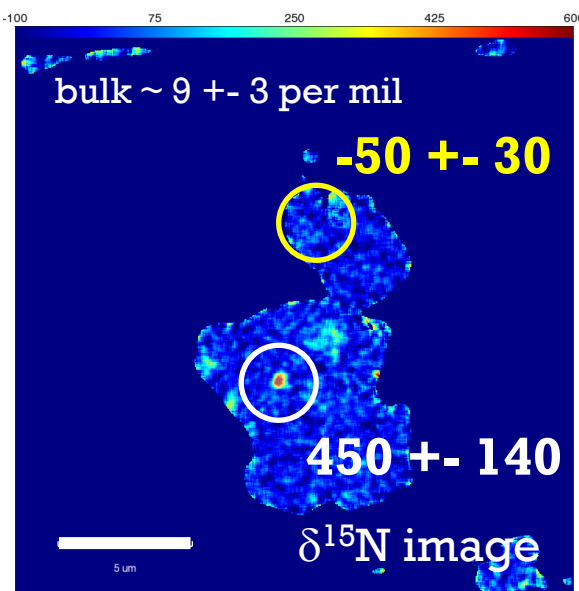
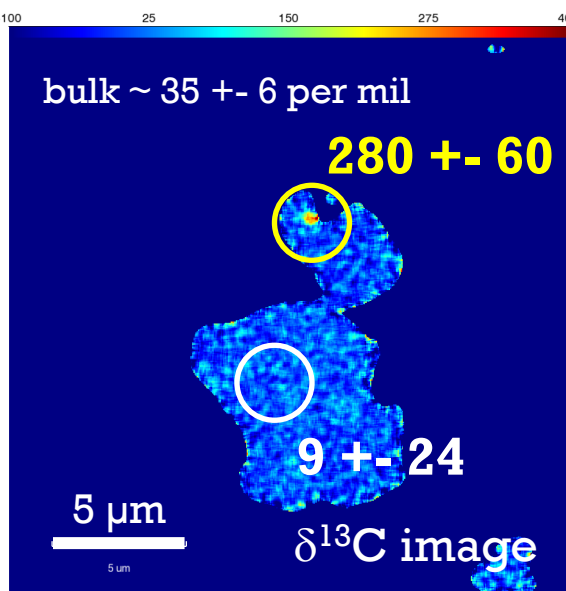
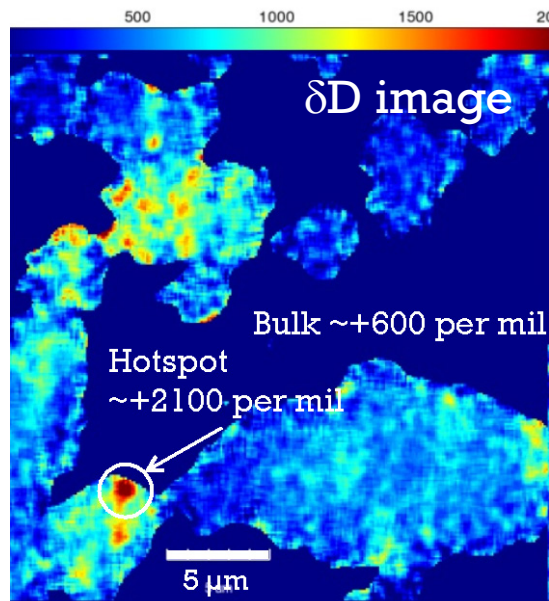
2. Outer regions of the protoplanetary disk

- large δD values have been predicted for gas-phase molecules.

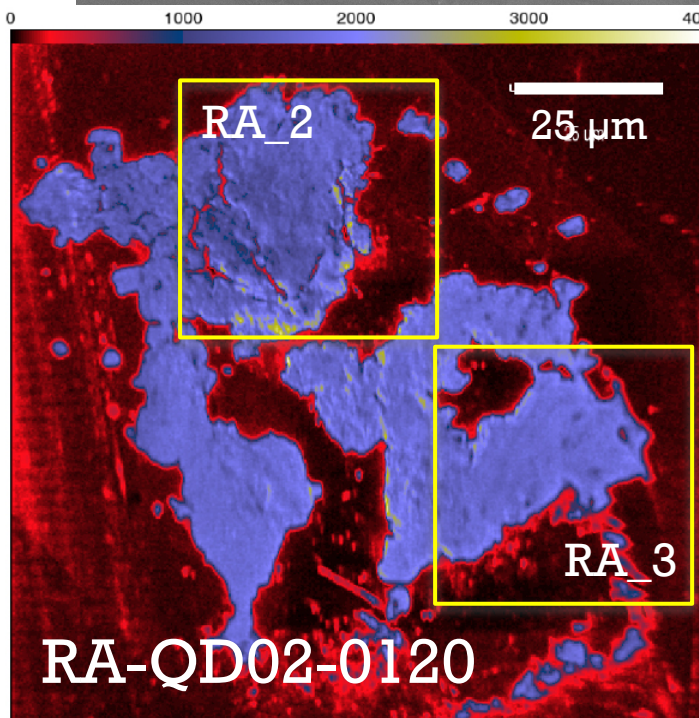
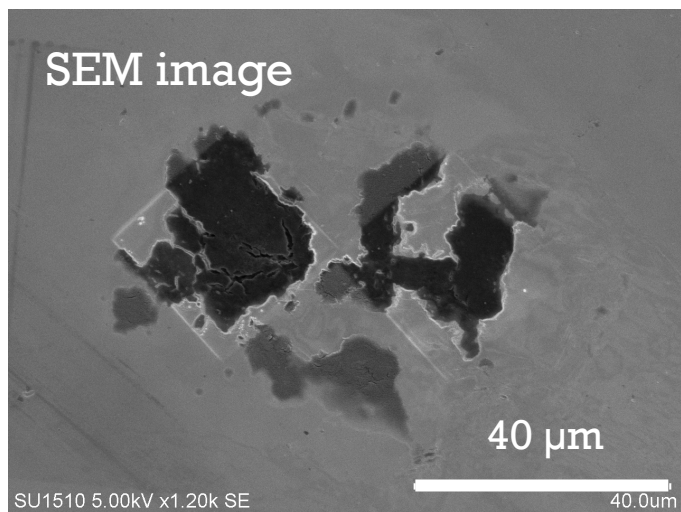
(e.g., Busemann et al., 2006)



Meteorites, IDPs, and comets are D-rich, preserving some presolar organic compounds

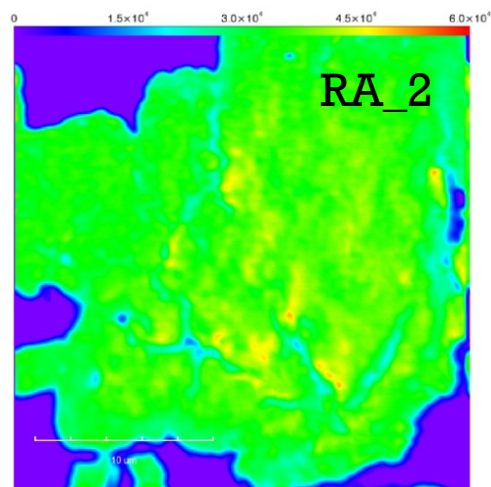


Hayabusa Category 3 organic sample

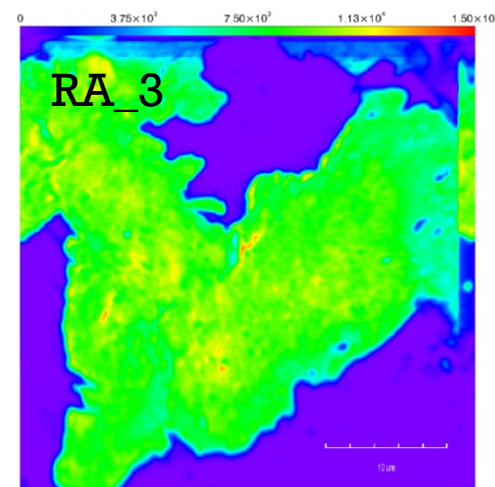


Entire image (^{12}C ^{14}N)

	$\delta^{13}\text{C}$ (per mil)	$\delta^{15}\text{N}$ (per mil)	δD (per mil)
RA_2	-19 \pm 8	2 \pm 2	81 \pm 54
RA_3	-6 \pm 8	0 \pm 2	29 \pm 47



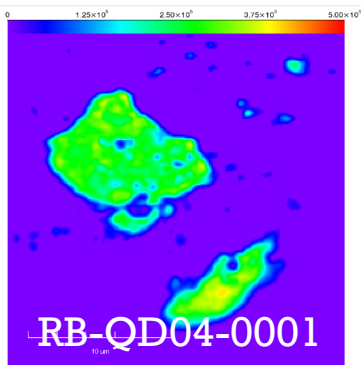
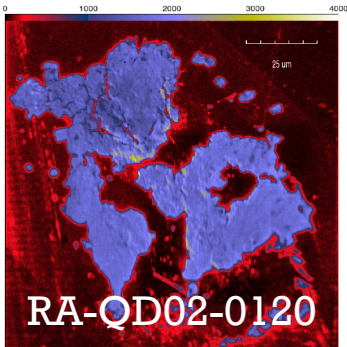
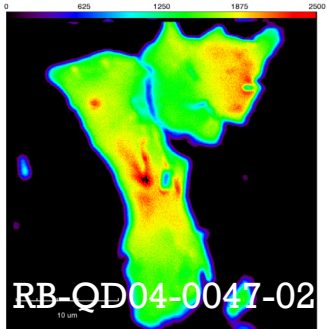
^{12}C ^{14}N image



Field of view = 25 μm



H and C isotope diagram



- RB-QD04-0047-02
- RA-QD02-0120
- RB-QD04-0001

δD (per mil)

Terrestrial

Extraterrestrial

$\delta^{13}C$ (per mil)

Sephton & Botta (2005)



Conclusions:



- Sample handling system for minerals, organic materials and microbes has been developed.
- Ion imaging by SIMS is a powerful tool to visualize distributions of isotopes and elements for variety of samples, and has a potential to allow the study of processes at the sub-micron level.
- Ultra high spatial resolution & quantitative isotope/elemental images will open up “***new insight***” for Astrobiology.



High sensitivity
and high
spatial
resolution SIMS

1. NanoSIMS
2. IMS1280HR
3. IMS6F

**ADVANCED MICROANALYSIS
RESEARCH LABORATORY
AT JAMSTEC KOCHI**

Ultra high
resolution TEM

1. High resolution TEM
2. FE-SEM
3. EELS, EDS, EBSD

Nano scale
sample
preparation
system

1. FIBs
2. Micro sampling system



Next Step: Cryo-linkage system will be added.