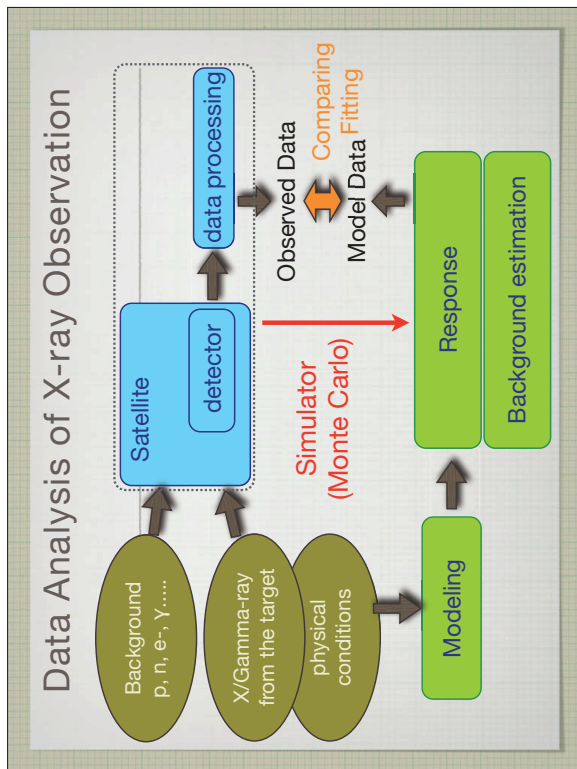
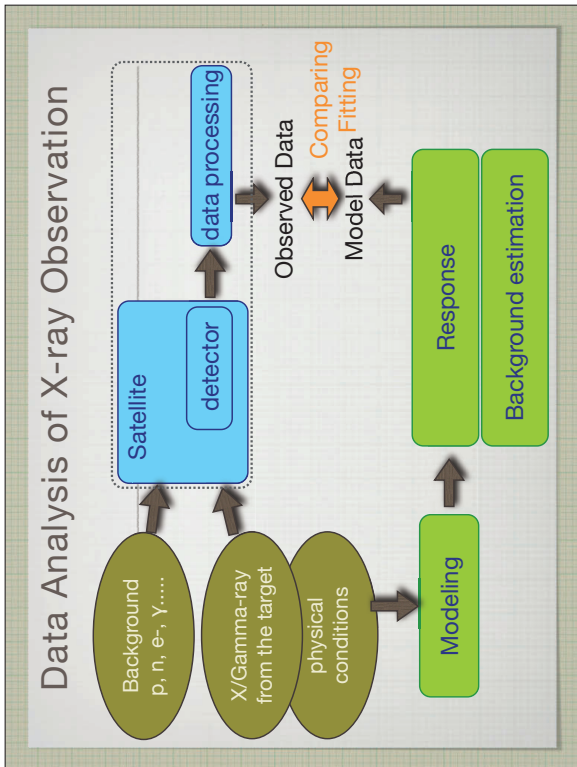


The X-ray simulators for Suzaku/NeXT

Shin Watanabe (ISAS/JAXA)

5th Geant4 Space Users' Workshop 2008
@University of Tokyo (13-15 Feb. 2008)



Suzaku(Astro-E2)




Suzaku: the 5th Japanese X-ray astronomy satellite

Launched on July 10, 2005

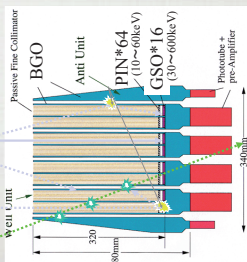
- XIS(X-ray CCD)
- HXD(hard X-ray detector)

Suzaku/HXD

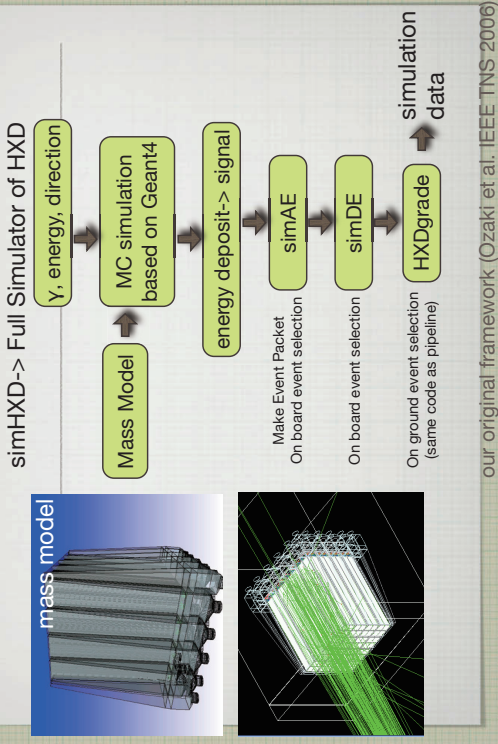
HXD:
 Si-PIN [2mm thick](10—70 keV)
 GSO [5mm thick](40—600keV)

Shield BGO well + Fine Collimator
 -> narrow FOV
 -> Low Background
 -> High Sensitivity

Complex Response
 Anti Coincidence, GSO-BGO phoswitch



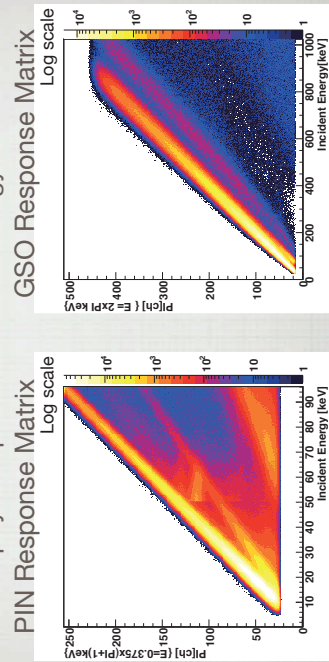
simHXD



HXD Response

HXD(PIN & GSO) responses constructed with simHXD distributed to world wide Suzaku users

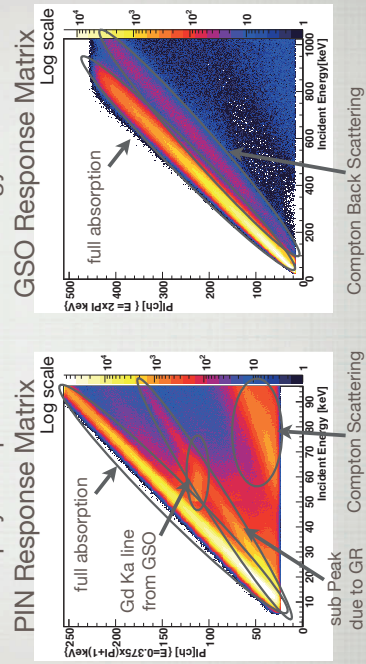
physical process: G4LowEnergyEM



HXD Response

HXD(PIN & GSO) responses constructed with simHXD distributed to world wide Suzaku users

physical process: G4LowEnergyEM

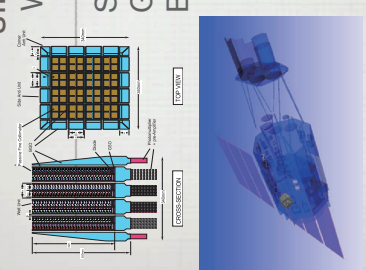


simHXD WAM

WAM: Wide All-sky Monitor
20 Large Volume BGO
Shield
GRB Detection & Spectroscopy
Bright Source Monitor
50 keV – 5 MeV

Absorption and scattering of solid Ne dewar, panels, solar array panel,


Response <- Constructed GRB by GRB
~70 GRBs have been analyzed with the simulator.



NeXT

New exploration X-ray Telescope [NeXT] (2013)
At Present, Pre-Project phase
We have constructed some prototype MC simulators.
Using prototype responses -> Science Topic Detector Design

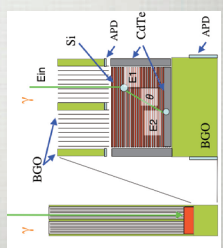
In Suzaku case, after FM design almost finished, Geant4 use started.
We plan to use Geant4 simulations for FM design of all science instruments



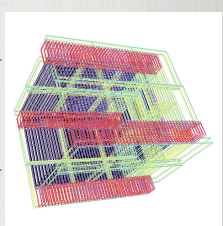
NeXT/SGD

SGD: Soft Gamma-ray Detector [50(10)–600 keV]
advance the concept of HXD (Low background)
Compton camera inside a well-type shield
using Compton reconstruction-> Background reduction
Compton camera -> Si-Pad and CdTe-Pad
for Gamma-ray response: G4LowEnergyEM+G4LECS(Kippen's talk)

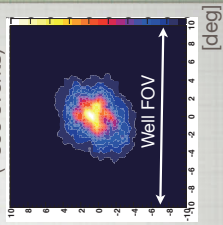
Concept of SGD



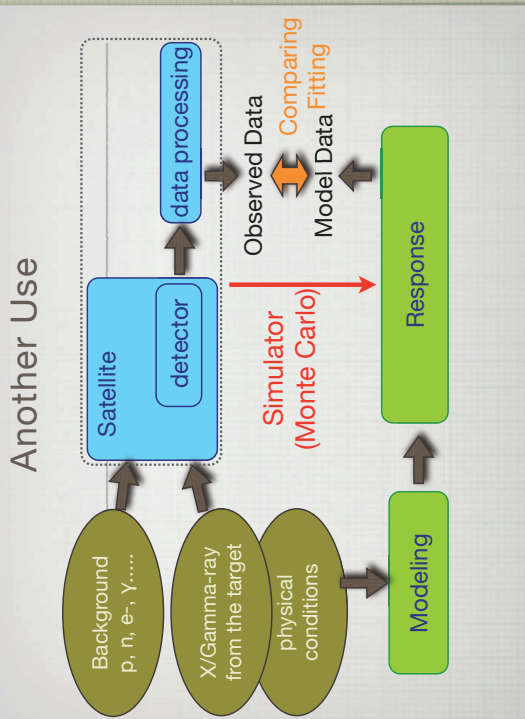
Mass Model of a Compton Camera part



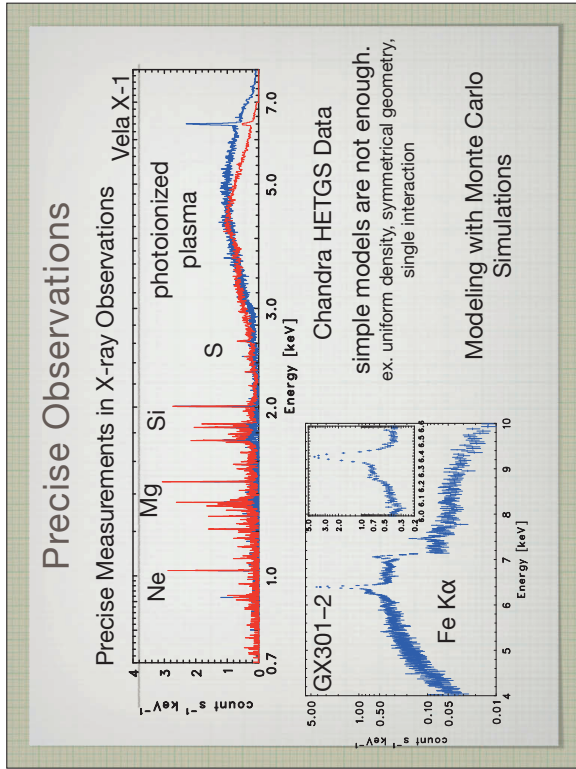
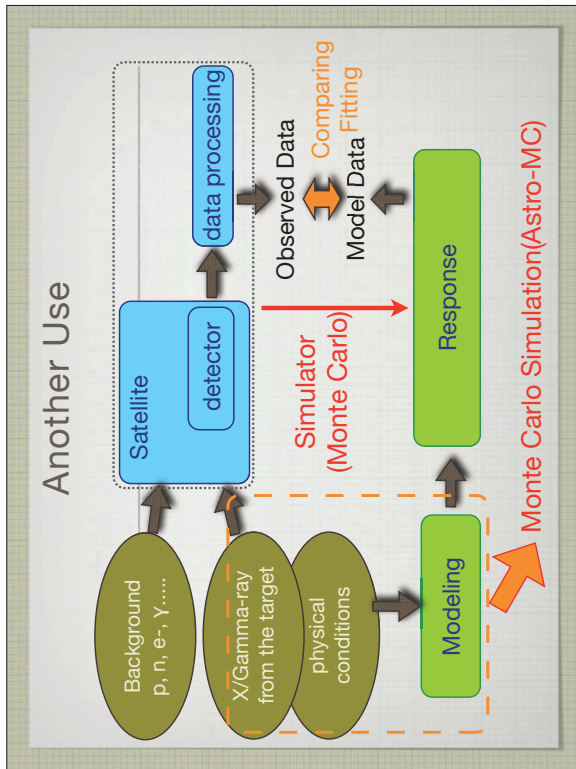
511 keV image (~500 events)



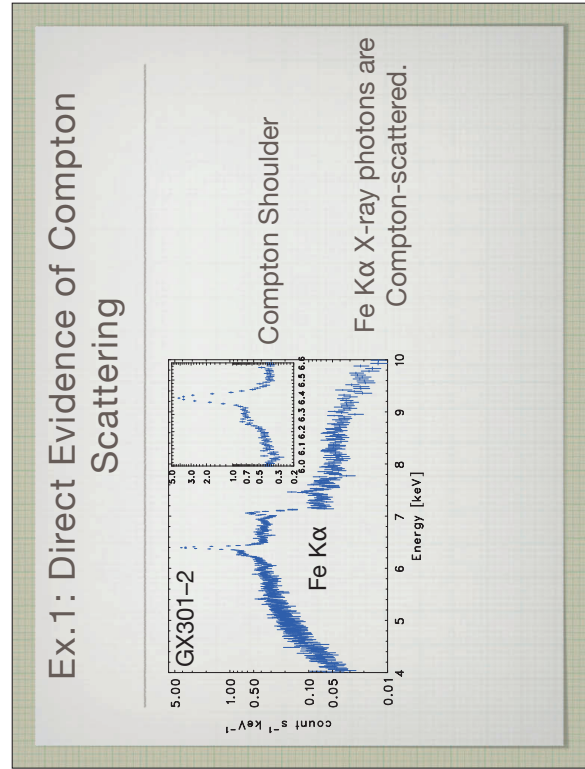
Another Use



Background p, n, e-, γ....
X/Gamma-ray from the target
physical conditions
Modeling
Simulator (Monte Carlo)
Satellite detector
data processing
Observed Data
Model Data Fitting
Response



- ### Geant4 and Astro-MC
- Issues
- Large Scale (10^{10}cm -- $> 10^{20}\text{cm}$)
 - no need to take care
 - scalable (geometry, physical process)
 - Physical Process
 - In some cases, need new physical processes (G4 physical processes \neq physical condition in space)
 - flexible



MC for a Dense Cloud

S. Watanabe et al. 2003 ApJL 597 pp. L37-L40

An X-ray source (neutron star) surrounded by a cloud
Compton Scattering process
H atoms are fully ionized and electrons have a temperature.

Simulation Results of Fe $K\alpha_1$, $K\alpha_2$

$NH = (1.20 \pm 0.35 - 0.13) \times 10^{24} \text{ cm}^{-2}$
 $kTe < 3.4 \text{ eV}$

Whole spectrum

MC Model and Observed Data

Fe absorption Edge
Second 20 msec
Fe $K\beta$ Line

An X-ray source surrounded by a cloud

reproduce the whole observed spectrum

Ex.2: Many Emission Lines from Plasma

Vela X-1 photoionized plasma

The ionization degree changes continuously

Physical Process Code for X-rays in Photoionized Plasma

New Physical Processes S. Watanabe et al. 2006 ApJ 651 pp.421-437

Photoionized Materials (C, N, O, Ne, Mg, Si, S, Ar, Ca, Fe)
H-like, He-like \rightarrow Photoionization (Recombination), Photoexcitation (Cascade)
Li-like to neutral \rightarrow Photoabsorption (crosssection and edge energy depend on the ionization stages.)
Compton Scattering
Doppler shift (Stellar Wind [100-1000 km/s])

Si He-like $kTe=5\text{eV}$
 10^{14} cm^{-2}
 10^{15} cm^{-2}
 10^{16} cm^{-2}
 10^{17} cm^{-2}

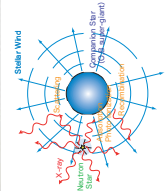
500 km/s

Si Cloud (H-like 50%, He-like 50%)

X-ray simulation in Photoionized Plasma

S. Watanabe et al. 2006 ApJ 651 pp.421-437

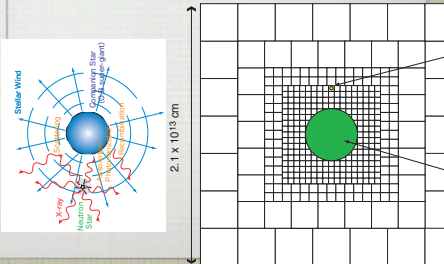
apply to Vela X-1 (neutron star
high mass X-ray binary)



X-ray simulation in Photoionized Plasma

S. Watanabe et al. 2006 ApJ 651 pp.421-437

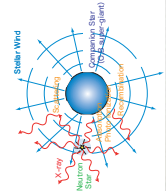
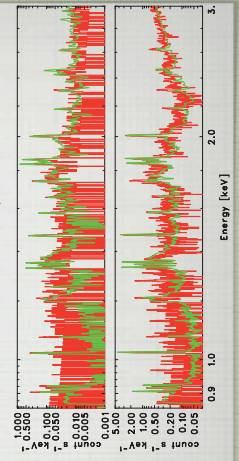
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X-ray simulation in Photoionized Plasma

S. Watanabe et al. 2006 ApJ 651 pp.421-437

apply to Vela X-1 (neutron star
high mass X-ray binary)

Summary

- Geant4 use for Suzaku (HXD Response, HXD Background, HXD-WAM)
- Geant4 use for NeXT and plan
- Soft Gamma-ray Detector case
- Astro-MC and examples of application (HIMXB X-ray spectroscopy)