

# The Estimation of Spatial Resolution of the Lunar Gamma-Ray Spectrometer aboard SELENE (KAGUYA) by Geant4

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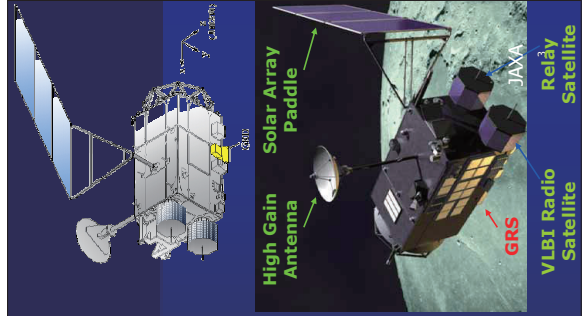
## Outline

- SELENE mission & Gamma ray spectrometer (GRS) onboard SELENE (KAGUYA)
- Spatial resolution of GRS
  - Importance of the spatial response function
  - Studies by using Geant4
    1. Emission angle of gamma ray at lunar surface
    2. Spatial response function of SELENE-GRS

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## SELENE (KAGUYA)

- Lunar explorer
- SELENE was launched by H-IIA rocket on Sep. 14, 2007.
- 14 scientific instruments
- Moon's origin and evolution
- Altitude: 100 km
- Polar orbit (~90 deg.)
- Attitude: 3 axes stabilized



## SELENE-GRS



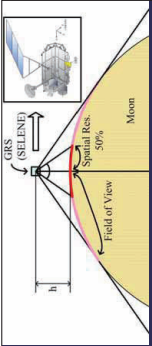
## Question

**Q. How large is the field of view of the GRS?**

A. GRS views the region from the nadir to the horizon, i.e. **any place** on the lunar surface. The detector is basically **omni-directional!**

→ But, an effective area observed by GRS (spatial resolution) is defined.

- The key is a **Spatial Response Function (SRF)**



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## What is a spatial response function?

- SRF is a set of probabilities to detect gamma rays as a function of positions (x, y) on the Lunar surface from which the detected gamma rays are emitted

$$I(\theta_1, \varphi_1) = \frac{1}{C_{\text{rate}}} \cdot \epsilon_{\text{int}}(\theta_3, \varphi_3) \cdot \omega(\theta_1, \varphi_1) \cdot f(\theta_2, \varphi_2, \theta_1, \varphi_1)$$

- **SRF is important.**
  - Spatial resolution of GRS
  - Incident direction dependences of the sensitivity
  - Deconvolution to make images

**Count rates at the altitude of spacecraft**

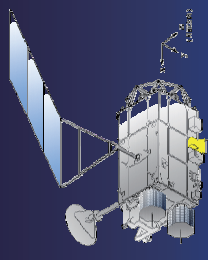
**SRF** deconvolution

**The flux of gamma ray at the surface**

3. Detection efficiency,  $\epsilon$

2. Solid Angle,  $\omega$

1. Distribution of Emission angle of gamma ray,  $f(\theta)$



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## Calculation by using Geant4 library

**Objectives**

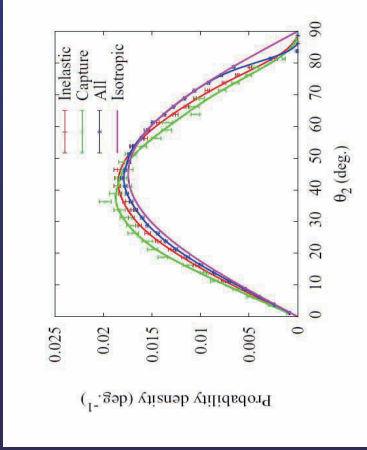
To understand the characteristics of the spatial response function ( $f_{\epsilon\omega}$ ) of SELENE-GRS

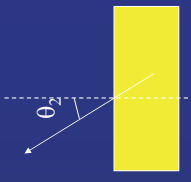
- 1.  $f(\theta)$ : Emission angle distribution of lunar gamma-ray**
  - Does the lunar surface emit gamma ray isotropically?
  - Emission angle will affect the SRF.
- 2. Calculation of the SRF of SELENE-GRS,  $f_{\epsilon\omega}$** 
  - Incident direction dependence of the sensitivity?
  - Spatial resolution of SELENE-GRS?

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## 1. Is the gamma-ray emission isotropic or not?, $f(\theta)$ ?

- The gamma rays are not emitted isotropically.
  - The gamma-rays tend to go upward.

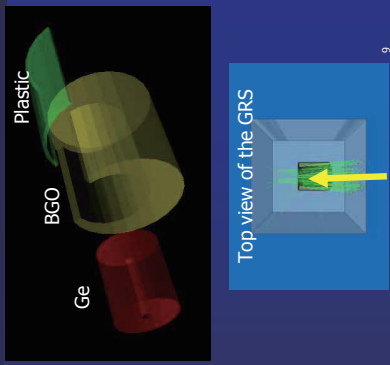




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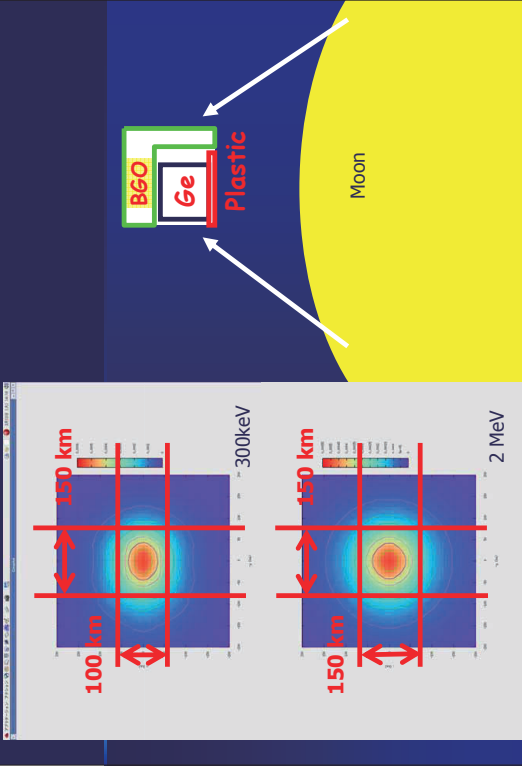
## 2. Calculation of spatial response function, $f \cdot \epsilon \cdot \omega$

- $f(\theta)$  <- isotropic
- $\epsilon \cdot \omega$  <- Geant4.8.0.p01
  - Detector
  - Ge+BGO+Plastic
  - Beam
    - gamma (0.1~10 MeV)
    - Beam dia. (~ 50 cm)
    - Incident directions (~3600)
      - polar angle,  $0 \sim 72^\circ$
      - azimuthal angle,  $0 \sim 180^\circ$

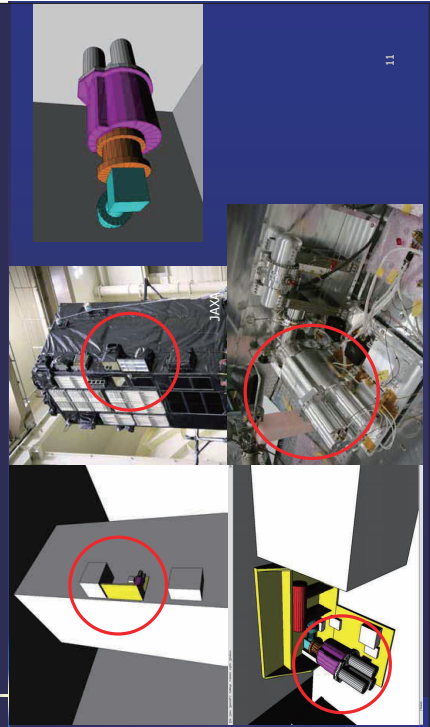


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## Contour plot of spatial response function ( $f \cdot \epsilon \cdot \omega$ )



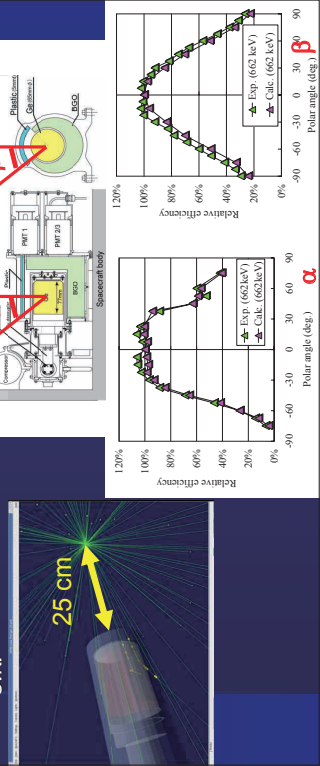
## Modeling of KAGUYA-GRS by Geant4



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## Model checking (exp. vs. calc.)

- Detection efficiency (JEEF standard method)
  - Exp. = 59.3 %
  - Calc. = 59 %
  - O.K.
- Detection efficiency (Pre-flight test)



## Summary

- Gamma-ray remote sensing is a powerful method
  - Global mapping of various elements (K, Th, U, Si, Fe...)
- The spatial response function is very important for GRS.
  - Spatial resolution
  - Incident direction dependences of the sensitivity
  - Deconvolution
- Calculation of SRF of SELENE-GRS was made by using Geant4 library
  - It calculates well
    - the anisotropy of sensitivity (SRF) of the SELENE-GRS,
    - energy dependence of SRF, and
    - the spatial resolution (150 km for 2 MeV)
- Current status
  - Checking the detailed model of SELENE-GRS made by Geant4 library